

CLASSICALIZATION AND QUANTIZATION OF TACHYON-LIKE MATTER ON (NON)ARCHIMEDEAN SPACES

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We consider a class of tachyon-like potentials, inspired by string theory, D-brane dynamics and cosmology in the context of classical and quantum mechanics. Motivated by the trans-Planckian problem in the very early stage of cosmological evolution of the Universe, we consider the theoretical role of DBI-type tachyon scalar field, defined over the field of real as well as p -adic numbers, i.e. archimedean and nonarchimedean spaces. To simplify the equation of motion for the scalar field, canonical transformations are defined and engaged. The corresponding quantum propagators in the Feynman path integral approach on real and nonarchimedean spaces are calculated and discussed, as are possibilities for a quantum adelic generalization and its application.

Key words: tachyons, quantum dynamics, DBI scalar field, inflation, quantum cosmology, nonarchimedean spaces.

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1. INTRODUCTION

According to the theory of cosmological inflation, the inhomogeneities in our Universe have a quantum-mechanical origin. This scenario is phenomenologically very appealing as it solves the problems of the standard hot big bang model and in a natural way explains the spectrum of cosmological perturbations [1].

In addition, one of the main goals of quantum cosmology is to find a quantum (gravitational) state which contributes to a consistent description of inflationary period of the Universe evolution. When accomplished, we could make predictions which should be more consistent with observational data.

According to the inflationary scenario, the Hubble radius was (almost) constant during inflation, the wavelength of a mode of astrophysical interest was much smaller than the Hubble scale at the beginning of inflation. Moreover, the modes are initially sub-Planckian, not only sub-Hubble, their wavelength is smaller than the Planck length. In this case the physics we are familiar with is not applicable any more [2, 3]. That leads us to idea to consider physical models on noncommutative and non-Archimedean spaces [4–6].

The general conclusion of various investigations in this direction of research is the observation that in the sub-Planckian (or trans-Planckian) regime the precise description of the evolution of the fluctuations is unclear. Various proposals are introduced in order to better understand the picture in the trans-Planckian regime, such as modification of dispersion relation by the trans-Planckian effects [3], modification of uncertainty relation [7], or the introduction of a new fundamental scale [8].

In keeping with this direction, we consider the DBI-type tachyon scalar field theory [9] in cosmological context as an effective field theory (taken from the string theory) which describes rolling tachyons [10]. This effective field theory has already been studied in the context of tachyonic inflationary cosmology, where, e.g. one can get kinematically driven inflation, the so-called "k-inflation" [11, 12]. Although it is argued that tachyonic inflation (on real numbers-spaces!) can not be responsible for the last 60 e-folds of inflation, it might be possible that tachyonic inflation is responsible for an earlier stage of inflation, at and "around" the Planck scale, which may be important for the resolution of homogeneity, flatness and isotropy problems [13]. For interesting ideas of p -adic inflation and p -adic rolling tachyons in a string theory context, see [14] and [15].

Here we will discuss the idea of using p -adic numbers (p stands for a prime number) in investigation of trans-Planckian effects in cosmology. Use of p -adic analysis is motivated and based on the simple fact that results of experimental and observational measurements always give some rational numbers. So, the field $R = Q_\infty$ of real numbers and the field Q_p of p -adic numbers are naturally put on the same footing, both are built by completion of the field of rational numbers with respect to the corresponding norms [16]. p -Adic norm is a non-Archimedean (ultrametric) one [17].

p -Adic numbers in cosmology are usually connected with quantum cosmology. Since quantum cosmology is used to describe the evolution of the universe at a very early stage, i.e. it is related to the Planck scale phenomena, then it is reasonable to consider various geometries, in particular the non-Archimedean one, which is closely connected to p -adic numbers and their application [4, 18].

p -Adic quantum mechanics has been developed in two different ways. Here we will use the one where the wave function is a complex valued function of the p -adic variable, at least for an adelic approach (simultaneous treatment of real and p -adic numbers) to quantum physics. This formulation is the preferred one [19] and has a natural, quantum, interpretation.

In order to explore theoretical models describing the early stage of cosmological evolution of the Universe (e.g. period of inflation) it is sufficient to start with spatially homogenous scalar fields [13, 20–24]. Cosmological perturbations are then introduced for the spatially homogenous scalar field and spatially homogenous background metric. As has already been noted, we will consider the DBI-type tachyon

scalar field, following Sen's proposals and conjectures [9], concentrating on the equation of motion for the field. Because this equation is not so simple even in the case of a spatially homogenous scalar field in the flat space-time background, we will try to deal with it imposing a canonical transformation treatment. This "transition" from field theory to a classical mechanical model [23, 25] we will call here "classicalization". Although it is an approach different from the idea recently proposed by Dvali and others [26], there are interesting similarities in equation of motions of DBI Lagrangians [27].

We will show that for certain choices of generating function of canonical transformation the equation of motion simplify significantly. At this point one can involve Lagrangian of the standard type which corresponds to the same equation of motion and which is locally equivalent [28, 29]. It allows us to quantize the simplified (toy) model, obtain quantum propagator and vacuum states in p -adic and adelic framework, as well as discuss their basic physical consequences.

This paper proceeds as follows. In Section 2 we will briefly review tachyonic cosmology, after which Section 3 deals with canonical transformation. In Section 4 certain choices of generating function of canonical transformation are introduced and presented through two examples. Section 5 deals with general remarks on classically equivalent Lagrangians ("classicalization"), while Section 6 presents quantization of the classical models, i.e. their real and p -adic consideration and a short review on a natural adelic generalization. We conclude in Section 7 by discussing obtained constraints on possible values of tachyon fields, time and constants appearing in two tachyonic models we consider.

2. TACHYONIC COSMOLOGY AND INFLATION

The Lagrangian we are dealing with - the DBI-type Lagrangian - is of non-standard type, it contains potential as a multiplicative factor and a term with derivatives ("kinetic" term) inside the square root [9, 30]

$$\mathcal{L}_{tach} = \mathcal{L}(T, \partial_\mu T) = -V(T) \sqrt{1 + (\partial T)^2}, \quad (1)$$

where T is tachyonic scalar field, $V(T)$ - tachyonic potential, $(\partial T)^2 = g_{\mu\nu} \partial^\mu T \partial^\nu T$ and $g_{\mu\nu}$ - components of the metric tensor, with "mostly positive signature". The tachyon potential $V(T)$ has a positive maximum at $T = 0$ and a minimum at T_0 with $V(T_0) = 0$ (T_0 can be either finite or infinite). General expression of the energy-momentum tensor

$$T_{\mu\nu} = \mathcal{L}_{tach} g_{\mu\nu} + V(T) \frac{\partial_\mu T \partial_\nu T}{\sqrt{1 + (\partial T)^2}}, \quad (2)$$

written in the form of the energy-momentum tensor for the ideal fluid defines pressure and energy density of a fluid described by the tachyonic scalar field

$$\rho = \frac{V(T)}{\sqrt{1 + (\partial T)^2}}, \quad (3)$$

$$p = \mathcal{L}_{tach}. \quad (4)$$

The state parameter w is then

$$w = \frac{p}{\rho} = -1 - (\partial T)^2. \quad (5)$$

One can define the so called effective sound speed c_s (for the cosmological perturbations [12])

$$c_s^2 = \frac{\partial \mathcal{L}_{tach}}{\partial (\partial T)^2} \left(\frac{\partial \rho}{\partial (\partial T)^2} \right)^{-1}, \quad (6)$$

which takes into account "friction" effects and it is a generic feature of theories with nonstandard Lagrangians of this type.

Equation of motion in curved spacetime is [10, 20]

$$D_\mu \partial^\mu T - \frac{D_\mu \partial^\nu T}{1 - (\partial T)^2} \partial_\mu T \partial_\nu T - \frac{1}{V(T)} \frac{dV}{dT} = 0, \quad (7)$$

where D_μ is covariant derivative with respect to $g_{\mu\nu}$. For the spatially homogenous tachyon field in flat spacetime background the last equation is reduced to

$$\ddot{T}(t) - \frac{1}{V(T)} \frac{dV}{dT} \dot{T}^2(t) = -\frac{1}{V(T)} \frac{dV}{dT}. \quad (8)$$

In this case one can start with the Lagrangian

$$\mathcal{L}_{tach}(T, \dot{T}) = -V(T) \sqrt{1 - \dot{T}^2}, \quad (9)$$

from which equation of motion (8) is obtained. In the classical mechanical limit T can be replaced by x and considered as a position variable in one-dimensional space [23, 25]. The equation contains a term proportional to \dot{T}^2 , which is, in general, less convenient to deal with. Note that even in the case of Minkowski spacetime background, the equation of motion for tachyonic field contains term with \dot{T}^2 , which is not the case for theories with standard-type Lagrangians. Objects w and c_s in this case correspond to each other in a simple manner [31]

$$w = -1 + \dot{T}^2 = -c_s^2. \quad (10)$$

Again, this kind of correspondence between w and c_s is not the case for the standard-type Lagrangian theories, where the effective sound speed is constant and equal to one (in units where speed of light $c = 1$) [31].

3. CLASSICAL CANONICAL TRANSFORMATION

In the light of the above discussion on tachyonic dynamics on (non)-Archimedean spaces, a natural question arises: is it possible to simplify the equation of motion (8) and, if it is possible, in which cases could that be done [23, 24, 28, 29]? Further, is it possible to quantize the models, at least for some particular potentials? A new possible way is to try with canonical transformation(s), which will be introduced and explained in this Section.

A classical canonical transformation is a change of the phase space variables (T, P) to a new (\tilde{T}, \tilde{P}) , which preserves the Poisson bracket

$$\{T, P\}_{P.B.} = 1 = \{\tilde{T}, \tilde{P}\}_{P.B.}. \quad (11)$$

We will seek for unitary transformations of coordinate (field) T and conjugate momenta P at the classical level

$$T, P \mapsto \tilde{T}, \tilde{P}, \quad (12)$$

$$\mathcal{H}_{tach}(T, P) \mapsto \tilde{\mathcal{H}}_{tach}(\tilde{T}, \tilde{P}), \quad (13)$$

which also preserves form of Hamilton's equations

$$\dot{T} = \frac{\partial \mathcal{H}_{tach}(T, P)}{\partial P} \rightarrow \dot{\tilde{T}} = \frac{\partial \tilde{\mathcal{H}}_{tach}(\tilde{T}, \tilde{P})}{\partial \tilde{P}}, \quad (14)$$

$$\dot{P} = -\frac{\partial \mathcal{H}_{tach}(T, P)}{\partial T} \rightarrow \dot{\tilde{P}} = -\frac{\partial \tilde{\mathcal{H}}_{tach}(\tilde{T}, \tilde{P})}{\partial \tilde{T}}. \quad (15)$$

Note the conjugate momenta and (conserved) Hamiltonian are [10]:

$$P = \frac{\partial L_{tach}}{\partial \dot{T}} = \frac{\dot{T}}{\sqrt{1 - \dot{T}^2}} V(T), \quad (16)$$

$$\mathcal{H}_{tach}(T, P) = \sqrt{P^2 + V^2(T)}. \quad (17)$$

We will use a particular model for generating function G , which specifies point canonical transformation. It is constructed as a function of new field \tilde{T} and old momenta P

$$G(\tilde{T}, P) = -PF(\tilde{T}), \quad (18)$$

where $F(\tilde{T})$ is an arbitrary function of a new field. By definition, the old coordinate T and the new momenta \tilde{P} are uniquely defined

$$T = -\frac{\partial G}{\partial P} = F(\tilde{T}), \quad (19)$$

$$\tilde{P} = -\frac{\partial G}{\partial \tilde{T}} = P \frac{dF(\tilde{T})}{d\tilde{T}}. \quad (20)$$

At the same time the new coordinate and old momenta can be expressed as

$$\tilde{T} = F^{-1}(T), \quad (21)$$

$$P = \frac{1}{F'(\tilde{T})} \tilde{P}, \quad (22)$$

where $F^{-1}(T)$ is an inverse function of $F(\tilde{T})$ and F' denotes derivative with respect to \tilde{T} . It is easy to check that Jacobian of this canonical transformation is equal to one and the Poisson brackets are unchanged, as expected.

Hamilton's equations become

$$\dot{\tilde{T}} = \frac{1}{F'} \frac{\tilde{P}}{\sqrt{\tilde{P}^2 + F'^2 V^2}}, \quad (23)$$

$$\dot{\tilde{P}} = -\frac{1}{F'^2} \frac{1}{\sqrt{\tilde{P}^2 + F'^2 V^2}} [(F')^2 V \frac{dV(F)}{dF} - F'' P^2], \quad (24)$$

while the equation of motion transforms to

$$\ddot{\tilde{T}} + \left(\frac{F''}{F'} - F' \frac{d \ln V(F)}{dF} \right) \dot{\tilde{T}}^2 + \frac{1}{F'} \frac{d \ln V(F)}{dF} = 0. \quad (25)$$

Note that equation (25) still contains term quadratic with respect to time derivative of a new coordinate (field) $\dot{\tilde{T}}$ (like (8)). Let us stress that this procedure, at the classical level, is formally invariant with respect to the choice of the background number fields R or Q_p .

4. CHOICE OF $F(\tilde{T})$

Up to now function $F(\tilde{T})$ was an arbitrary one. If the function $1/V(T)$ is integrable, function $F(\tilde{T})$ can be defined in such a way that its inverse function $F^{-1}(T)$ is equal to

$$F^{-1}(T) = \int^T \frac{dT}{V(T)}, \quad (26)$$

where the lower limit of the integral can be chosen arbitrary. This particular choice enables us to simplify equation of motion (25) significantly. In particular, the second term in (25) vanishes in this case, because the expression in parenthesis is identically equal to zero

$$\frac{F''}{F'} - F' \frac{d \ln V(F)}{dF} = 0. \quad (27)$$

So, using (26), the equation of motion (25) is reduced to the form

$$\ddot{\tilde{T}} + \frac{1}{F'} \frac{d \ln V(F)}{dF} = 0. \quad (28)$$

Note that equation (28) now stands for the system without term quadratic with respect to $\dot{\tilde{T}}$ (unlike (8) and (25)). The term disappeared after we imposed suitable chosen canonical transformation, i.e. choice (26).

It is very important to emphasize that equation of motion (25) in some cases can be obtained from the standard type Lagrangians [23, 24, 32]. Moreover, equation (28) in some cases can be obtained from the standard type Lagrangians which are quadratic with respect to \tilde{T} and $\dot{\tilde{T}}$, a situation very suitable for the Feynman path integral approach for quantization. From several interesting tachyonic or tachyon-like potentials we have considered, we choose two of the most used class of potentials to study here in more detail: $V(T) = e^{-\alpha T}$ and $V(T) = 1/\cosh(\beta T)$ [22, 33].

4.1. EXAMPLE 1: EXPONENTIAL POTENTIAL

The case of exponential potential of a tachyonic field

$$V(T) = e^{-\alpha T}, \quad \alpha > 0 - const, \quad (29)$$

is well known and motivated by string theory [9]. The function $F^{-1}(T)$ becomes

$$F^{-1}(T) = \frac{1}{\alpha} e^{\alpha T}, \quad (30)$$

which means that function $F(\tilde{T})$, using (19), becomes

$$F(\tilde{T}) = \frac{1}{\alpha} \ln(\alpha \tilde{T}). \quad (31)$$

The full generating function (18)

$$G(\tilde{T}, P) = -PF(\tilde{T}) = -\frac{P}{\alpha} \ln(\alpha \tilde{T}), \quad (32)$$

reduces equation of motion to the well known form

$$\ddot{\tilde{T}} - \alpha^2 \tilde{T} = 0. \quad (33)$$

This equation of motion can be delivered from a quadratic Lagrangian

$$\mathcal{L}_{quad}(\tilde{T}, \dot{\tilde{T}}) = \frac{1}{2} \dot{\tilde{T}}^2 + \frac{1}{2} \alpha^2 \tilde{T}^2. \quad (34)$$

4.2. EXAMPLE 2: POTENTIAL $V(T) = 1/\cosh(\beta T)$

As in the previous case, with a similar motivation, we consider the potential

$$V(T) = \frac{1}{\cosh(\beta T)}, \quad \beta > 0 - const, \quad (35)$$

and the function $F^{-1}(T)$ becomes

$$F^{-1}(T) = \frac{1}{\beta} \sinh(\beta T), \quad (36)$$

what means that function $F(\tilde{T})$, using again (19), is

$$F(\tilde{T}) = \frac{1}{\beta} \operatorname{arcsinh}(\beta \tilde{T}). \quad (37)$$

The full generating function (18)

$$G(\tilde{T}, P) = -PF(\tilde{T}) = -\frac{P}{\beta} \operatorname{arcsinh}(\beta \tilde{T}), \quad (38)$$

reduces equation of motion again to the form

$$\ddot{\tilde{T}} - \beta^2 \tilde{T} = 0. \quad (39)$$

Here we find the same situation, this equation of motion can be delivered from a quadratic Lagrangian

$$\mathcal{L}_{quad}(\tilde{T}, \dot{\tilde{T}}) = \frac{1}{2} \dot{\tilde{T}}^2 + \frac{1}{2} \beta^2 \tilde{T}^2. \quad (40)$$

Let us stress again that this equivalence is a "local" one, based on (26). Keeping in mind that motivation for this study comes from inflation theory and should be applied to its very beginning (quantum origin and very short period of time), this simplification seems quite meaningful for further consideration.

5. EQUIVALENT LAGRANGIANS QUANTIZATION REVISITED

As we noticed in the previously considered simplified cases it is possible to pass from the non-standard (DBI) Lagrangian to the locally equivalent and "canonical" one. In both cases, we ended up with the quadratic Lagrangian, with the potential terms having the "wrong" sign, i.e. the form of an inverted harmonic oscillator Lagrangian [34].

Its repulsive ("wrong" sign), negative pressure-like and "antigravitational" effect mimics "dark energy" and offers a playground for formal speculation and consideration of its real, p -adic and adelic origin. This will be discussed elsewhere. On speculation of "p-adic dark energy" see, for instance, [35].

At the (real, i.e. non p -adic) quantum level, inverted oscillator system has an energy spectrum, varying from minus to plus infinity [34]. So, the state with the lowest energy corresponds to negative infinite energy, $E = -\infty$. The general solution of Schroedinger equation for the inverted oscillator can be presented as a linear combination of solutions with definite parity [36]

$$\Psi(\tilde{T}) = C\Psi_{even}(\tilde{T}) + D\Psi_{odd}(\tilde{T}), \quad (41)$$

where C and D are real constants and Ψ_{even} and Ψ_{odd} are expressed in terms of confluent hyperbolic functions.

By introducing "annihilation" and "creation" operators, as was done for the harmonic oscillator, one ended up with the so-called generalized eigenstates belonging to the complex energy eigenvalues. As it is known, the energy eigenvalue E can be a complex number for an unstable system for which the potential energy does not have a stable stationary point. That is the case here (see [37] and reference therein for the discussion about mathematical formulations of continuous spectrum or complex eigenvalues on real space).

One can ask for which (tachyon-like) potentials present in the DBI Lagrangian it is possible to get quadratic Lagrangian and with which sign, i.e. formally with an attractive or repulsive force?

Invoking classical canonical transformation, first of all to simplify equation of motion, we deliver an equation which can be obtained from the various standard and non-standard type Lagrangians. At the classical level, we can choose the most suitable Lagrangian for the system. However, at the quantum level, this is not so simple. As different Lagrangians can lead to different quantum systems.

Regarding the quantum level, the path integral formulation of quantum mechanics (and also other equivalent formulations) depends on a Lagrangian chosen to describe the classical system. The arbitrariness in this choice of suitable classical Lagrangian leads to the so-called quantization ambiguity.

Nevertheless, in the next Section we will discuss the quantum mechanical propagator on both real and p -adic spaces (Archimedean and non-Archimedean ones), constructed from the most suitable Lagrangian which gives the same equation of motion as the original one. This is related to the idea [38] in which functional integral formula for the quantum propagator could be defined starting from the given classical equation of motion and, hence, would not refer to the form of the chosen Lagrangian.

6. REAL, p -ADIC AND ADELIC CONSIDERATION

As noted in [15], it was shown that p -adic string resembles the bosonic string in such way that its ground state is a tachyon, whose unstable maximum presumably indicates the presence of a decaying brane, analogous to the unstable D25-brane of the open bosonic string theory [9].

A possibility to get successful inflation from rolling p -adic tachyons toward the bounded direction was shown in [14]. This is an opposite and very interesting appearance compared to the real case [13], because the tachyon potential is not flat enough to give a significant period of inflation. Due to nonlocality and the nonarchimedean character of p -adic space, p -adic string tachyon can roll slowly enough to give many e -foldings of inflation. Nevertheless, it is natural to expect that p -adic effects should be of some importance for the initial phase of inflation and its "very

first e -foldings". However, all similar considerations of (p -adic) tachyons have been undertaken in a classical manner and quantum dynamics remains to be studied.

Hence, we start our quantum consideration from Lagrangian (34) or (40), which are quadratic with respect to \tilde{T} and $\dot{\tilde{T}}$. We can write down the corresponding transition amplitudes (propagators) on real spaces and using real numbers [39],

$$\mathcal{K}_\infty(\tilde{T}_2, \tau; \tilde{T}_1, 0) = \sqrt{-\frac{1}{2\pi i \hbar} \frac{\partial^2 S_c}{\partial \tilde{T}_1 \partial \tilde{T}_2}} e^{i \frac{S_c}{\hbar}}, \quad (42)$$

where S_c is a classical action of the system, which is quadratic with respect to initial and final configuration \tilde{T}_1 and \tilde{T}_2 . Here, τ denotes elapsed time. It can be also written in the form [40]

$$\begin{aligned} \mathcal{K}_\infty(\tilde{T}_2, \tau; \tilde{T}_1, 0) &= \lambda_\infty \left(-\frac{1}{2\hbar} \frac{\partial^2 S_c}{\partial \tilde{T}_1 \partial \tilde{T}_2} \right) \left| \frac{1}{\hbar} \frac{\partial^2 S_c}{\partial \tilde{T}_1 \partial \tilde{T}_2} \right|_\infty^{1/2} \times \\ &\quad \chi_\infty \left(-\frac{1}{\hbar} S_c(\tilde{T}_2, \tau; \tilde{T}_1, 0) \right), \end{aligned} \quad (43)$$

where an arithmetic λ -function and additive character χ_∞ are defined as

$$\lambda_\infty(b) = e^{-\frac{i\pi}{4} \text{sgn}(b)}, \quad \chi_\infty(a) = e^{-2\pi i a}. \quad (44)$$

Let us discuss the transition amplitude in the p -adic case. It can be formally done by changing the number field, from $R = Q_\infty$ to Q_p . This means that we will deal with p -adic numbers and complex wave functions of p -adic argument [17].

The transition amplitude in p -adic case \mathcal{K}_p , for an action quadratic in \tilde{T}_1 and \tilde{T}_2 (we take $\hbar = 1$ for simplicity), as it was shown in [40] is

$$\mathcal{K}_p(\tilde{T}_2, \tau; \tilde{T}_1, 0) = \lambda_p \left(-\frac{1}{2} \frac{\partial^2 S_c}{\partial \tilde{T}_2 \partial \tilde{T}_1} \right) \left| \frac{\partial^2 S_c}{\partial \tilde{T}_2 \partial \tilde{T}_1} \right|_p^{1/2} \chi_p \left(-S_c(\tilde{T}_2, \tau; \tilde{T}_1, 0) \right), \quad (45)$$

where the p -adic additive character χ_p is defined as [17]

$$\chi_p(a) = e^{2\pi i \{a\}_p}, \quad (46)$$

$\{a\}_p$ is the fractional part of the p -adic number a , while λ_p is an arithmetic complex-valued function (here with a p -adic variable), with the following basic properties

$$\lambda_p(0) = 1, \quad \lambda_p(a^2 b) = \lambda_p(b), \quad |\lambda_p(b)|_\infty = 1, \quad (47)$$

$$\lambda_p(a) = 1, \quad |a|_p = p^{\text{ord}(a)} = p^{2\gamma}, \quad \gamma \in \mathbb{Z}. \quad (48)$$

In this way, the transition amplitude \mathcal{K}_p for the Lagrangians (34), (40) has the

form

$$\begin{aligned} \mathcal{K}_p(\tilde{T}_2, \tau; \tilde{T}_1, 0) = & \lambda_p \left(\frac{\gamma}{2\sinh(\gamma\tau)} \right) \left| \frac{\gamma}{\sinh(\gamma\tau)} \right|_p^{1/2} \\ & \times \chi_p \left(-\frac{\gamma}{2} \left((\tilde{T}_1^2 + \tilde{T}_2^2) \coth(\gamma\tau) - \frac{2\tilde{T}_1\tilde{T}_2}{\sinh(\gamma\tau)} \right) \right), \end{aligned} \quad (49)$$

where γ stands either for α regarding Lagrangian (34) or β regarding (40), respectively.

The necessary condition for the existence of a p -adic and an adelic model is to find a p -adic quantum-mechanical ground state in the form of Ω -function

$$\Omega(|\tilde{T}|_p) = \begin{cases} 1, & \text{if } |\tilde{T}|_p \leq 1 \\ 0, & \text{if } |\tilde{T}|_p > 1, \end{cases} \quad (50)$$

which is characteristic function on Z_p , the set of p -adic integers. A physical interpretation is that a system (particle) remains in its vacuum (ground) state as long as it is considered to be inside a "box" with a "natural" unity length (for example Planck length l_{pl}), etc. Note that Ω -function is a counterpart of the Gaussian $\exp(-\pi\tilde{T}^2)$ in the real case ($\tilde{T} \in R$), since it is invariant with respect to the Fourier transform [4].

Having in mind one of the basic properties of (p -adic) propagator and the corresponding unitary operator of evolution

$$\hat{U}(\tau)\Psi_p^{vac}(\tilde{T}_2) = \int_{Q_p} \mathcal{K}_p(\tilde{T}_2, \tau; \tilde{T}_1, 0) \Psi_p^{vac}(\tilde{T}_1) d\tilde{T}_1, \quad (51)$$

we get for the vacuum state $\Psi_p^{vac}(\tilde{T}) = \Omega(|\tilde{T}|_p)$

$$\int_{Z_p} \mathcal{K}_p(\tilde{T}_2, \tau; \tilde{T}_1, 0) d\tilde{T}_1 = \Omega(|\tilde{T}_2|_p), \quad (52)$$

where the interval of integration becomes restricted to Z_p [40].

The necessary conditions for the existence of ground states in the form of the characteristic Ω -function are "dictated" by the expression (52), from which the necessary restriction on the values of the parameters of the theory can be found [29]. For $p \neq 2$ the necessary conditions are

$$\Psi_p^{vac}(\tilde{T}) = \Omega(|\tilde{T}|_p), \quad \text{for } \begin{cases} |\tau|_p = 1, \\ |\tau|_p < 1, \end{cases} \quad |\gamma^2 \tilde{T}^2 \tau|_p \leq 1. \quad (53)$$

Note that in p -adic quantum mechanics there is degeneration of the vacuum state. There are other possibilities for the ground states wave function: "modified" Ω -function $\Omega(p^\nu |\tilde{T}|_p)$ and p -adic Dirac delta-function $\delta(p^\nu - |\tilde{T}|_p)$, $\nu \in Z$ [4]. They will give another set of conditions for the existence of ground state and restrictions for the parameters [40]. This consideration will be discussed elsewhere.

For the theory under consideration, the adelic wave function would be of the form

$$\Psi_{ad}(\tilde{T}_a) = \prod_v \Psi_v(\tilde{T}_v) = \Psi_\infty(\tilde{T}_\infty) \prod_{p \in M} \Psi_p(\tilde{T}_p) \prod_{p \notin M} \Omega(|\tilde{T}_p|_p), \quad (54)$$

where $v = (\infty, 2, 3, \dots, p, \dots)$, M is a finite set of primes p , while $\tilde{T}_\infty \in R$ and $\tilde{T}_p \in Q_p$ defines an adele \tilde{T}_a , i.e. a sequence of the form

$$\tilde{T}_a = (\tilde{T}_\infty, \tilde{T}_2, \tilde{T}_3, \dots, \tilde{T}_p, \dots), \quad (55)$$

where for all but finitely many p , $|\tilde{T}|_p \leq 1$. The corresponding adelic transition amplitude (quantum propagator) as a non-trivial product of (43) and (49) for all p is [4]

$$\begin{aligned} \mathcal{K}_{ad}(\tilde{T}_2, \tau; \tilde{T}_1, 0) &= \prod_v \mathcal{K}_v(\tilde{T}_2, \tau; \tilde{T}_1, 0) = \\ &= \mathcal{K}_\infty(\tilde{T}_2, \tau; \tilde{T}_1, 0) \times \prod_p \mathcal{K}_p(\tilde{T}_2, \tau; \tilde{T}_1, 0). \end{aligned} \quad (56)$$

Note that $\Psi_\infty(\tilde{T}_\infty)$ is the corresponding real (counterpart of the) wave functions of the theory, and $\Psi_p(\tilde{T}_p)$ are the p -adic wave function (again, for all but finitely many p , $\Psi_p(\tilde{T}_p) = \Omega(|\tilde{T}_p|_p)$). In case of a "true" adelic vacuum state $\Psi_p(\tilde{T}_p) = \Omega(|\tilde{T}_p|_p)$ for all p .

7. CONCLUSION

We considered the DBI-type tachyon scalar field theory, inspired by a common belief that tachyon field can be used in the frame of cosmology, in particular in the initial phase of inflation. We discussed the idea to extend a standard approach using p -adic numbers and ultrametric geometry and spaces or, more generally, non-Archimedean ones. This, at the very least, bearing in mind that all results of experimental and observational measurements are always some rational numbers, justifies interest in this generalization [16, 17].

We presented the possibility of passing from the non-standard (DBI) Lagrangian to the standard and very familiar one (i.e. Lagrangian for inverted harmonic oscillator). In this paper we presented an original canonical transformation of a particular form, suitable for transformation of a class of relevant tachyonic potentials. We calculated the corresponding quantum propagators using the Feynman path integral approach in two special cases (exponential and *cosh* potential) and presented their form on real, p -adic and adelic spaces. A set of conditions for the existence of vacuum states in p -adic and adelic cases, which restricts the allowed domain, for time, tachyonic field ("distance" in "one-dimensional" classical mechanical limit of field theory) and parameters present in considered potentials were obtained. For instance,

one can speculate that the tachyonic scalar field exists in an adelic vacuum state as long as $|\tau|_p < 1$ (for instance, $t_{pl} = 1$, or some other natural unity of time) and predict $|\gamma^2 \tilde{T}^2 \tau|_p \leq 1$. It is obvious that (53) gives a very interesting constraint on values of tachyon field T , elapsed time τ and constant parameters of the theory γ (i.e. α and β). It can be used for calculation and prediction of values of (tuning parameter) constants γ , which determine the shape of potentials (29) and (35). It can also be connected to allowed values of position of "classical particle" counterpart and tachyonic field.

The theory of p -adic inflation is not yet "phenomenological" enough to predict concrete values of these constants (γ , \tilde{T} , τ , second line of (53)) and corresponding parameters of inflation. Thoughts on that, including solution of Friedmann-like equations, will be presented elsewhere [41]. We find that our approach based on canonical transformation can be applied to inverse power tachyon potentials [42]. Nearly quadratic systems with a term which could be studied as a perturbation of a quadratic system is also a very interesting task.

We find quite interesting that the tachyonic system provides rich and fruitful cosmological scenarios for inflation, as well as for dark components, to possess non-trivial quantum dynamics and deserve further attention and consideration. We would like to underline that our approach can be useful in a complete "real" consideration, when all p -adic and adelic extensions are neglected.

Finally, we mention one of the open problems: how to connect wave function of the Universe [4] with tachyon quantum effects and wave function (54), i.e. Ψ_∞ , Ψ_p and Ψ_{ad} .

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